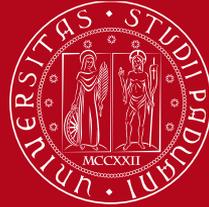


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In-flight calibration and performance verification of the Metis and STC instruments

Chiara Casini - 36th Cycle

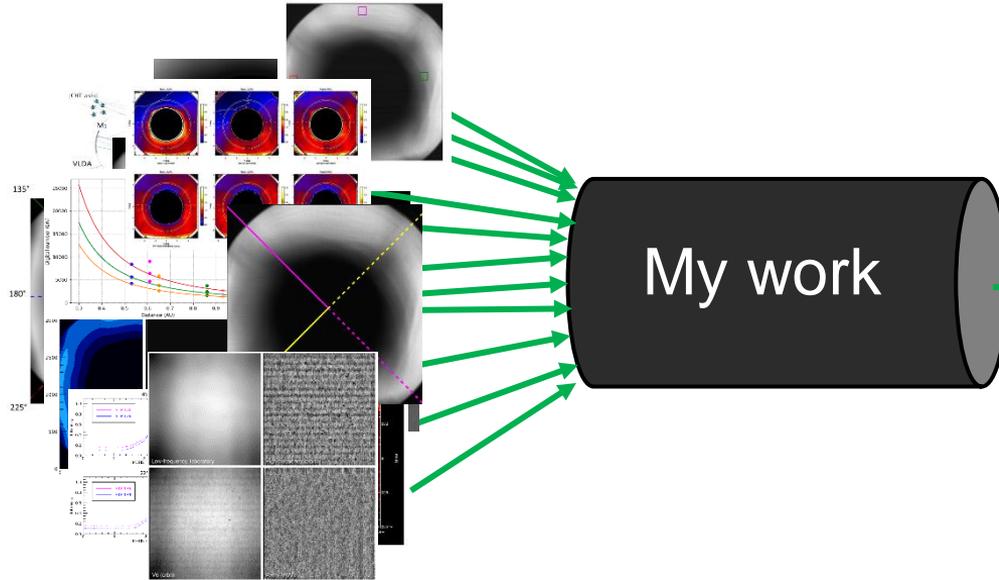
MEETING FOR THE ADMISSION TO THE THIRD YEAR

05-06/09/2022

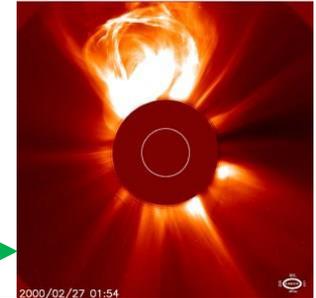
~~SCIENCE???~~



~~NO~~

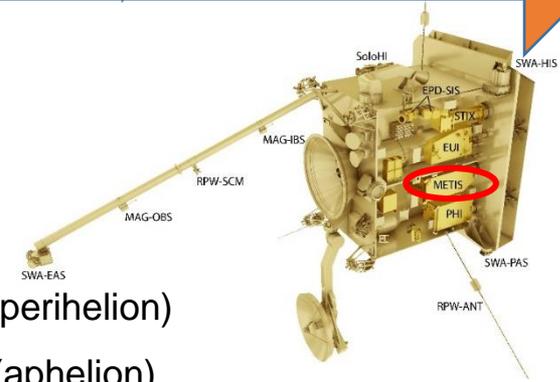
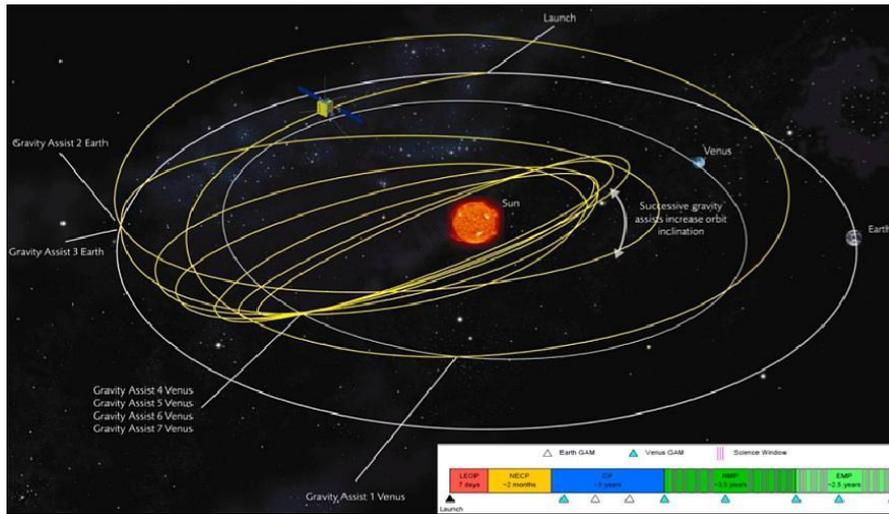


SCIENCE???



YES

My PhD activity serves to acquire a deep and detailed knowledge of the Metis and STC instruments in one of the most critical and important phase of a space mission: it's essential to obtain scientific useful images.



ORBIT:

0.28 – 0.32 AU (perihelion)

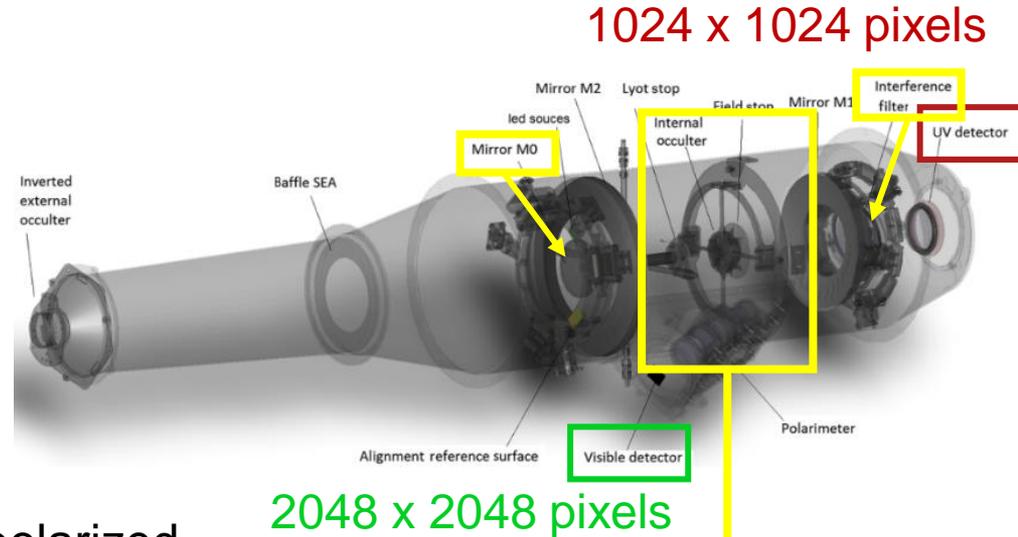
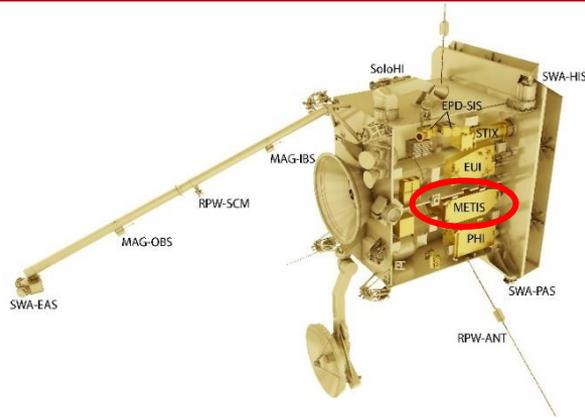
0.74 -- 0.91 AU (aphelion)

Out-of-ecliptic view:

Multiple gravity assists with Venus to increase inclination out of the ecliptic to :

>24° (nominal mission);

>34° (extended mission)



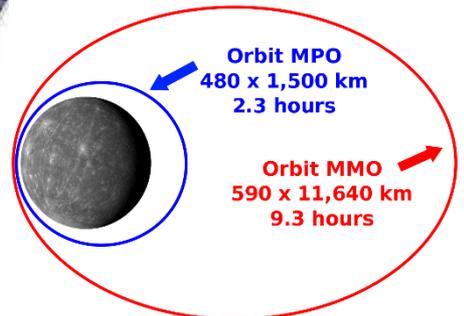
Full Imaging of the corona ($1.7 - 9 R_{\odot}$):

- UV (121.6 ± 10 nm)
- visible light (580-640nm) in total and polarized brightness

Spatial resolution and detector exposure time can change based on the science goal.

The Internal Occulter (IO) is extremely important to minimize the straylight.

Launch BepiColombo



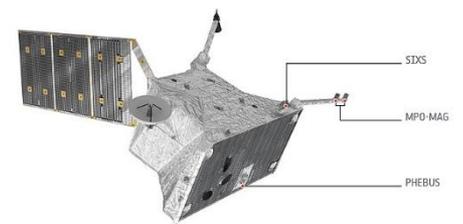
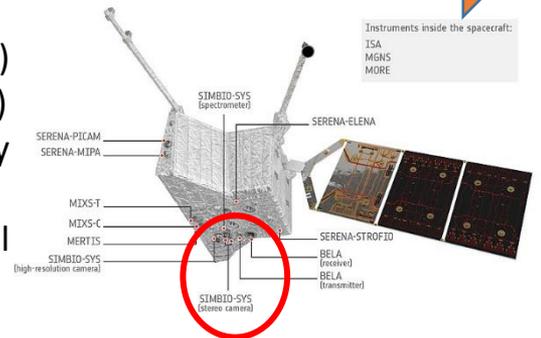
BepiColombo is based on two modules:

- **Mercury Magnetospheric Orbiter (MMO)** realized by the JAXA (Japanese space agency)
- **Mercury Planetary Orbiter (MPO)** realized by the ESA.

MPO study: surface, exosphere and internal composition of the planet.

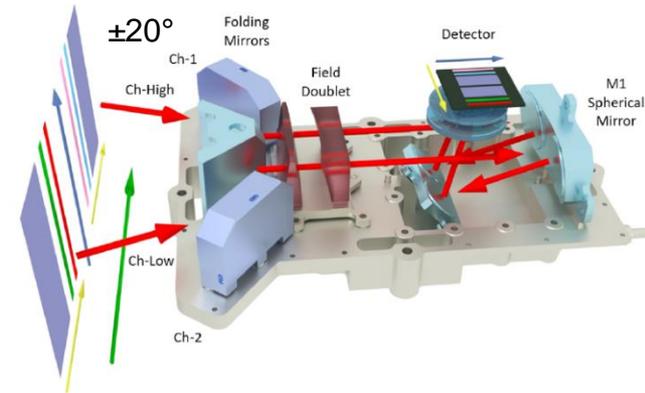
- ✓ Periherm and apoherm [480 km, 1500 km].
- ✓ Orbital period of 2.3 hours.

MPO accommodates 11 instruments including SIMBIO-SYS where STC is included.





STC main scientific objective is the global mapping of the entire surface of Mercury in 3D.



The STC camera is able to reach the goal thanks to the two sub-channels: High (H) and Low (L).

- a front unit, which consists of two independent fore-optics modules, one for each subchannel;
- a common telescope unit (off-axis modified Schmidt)

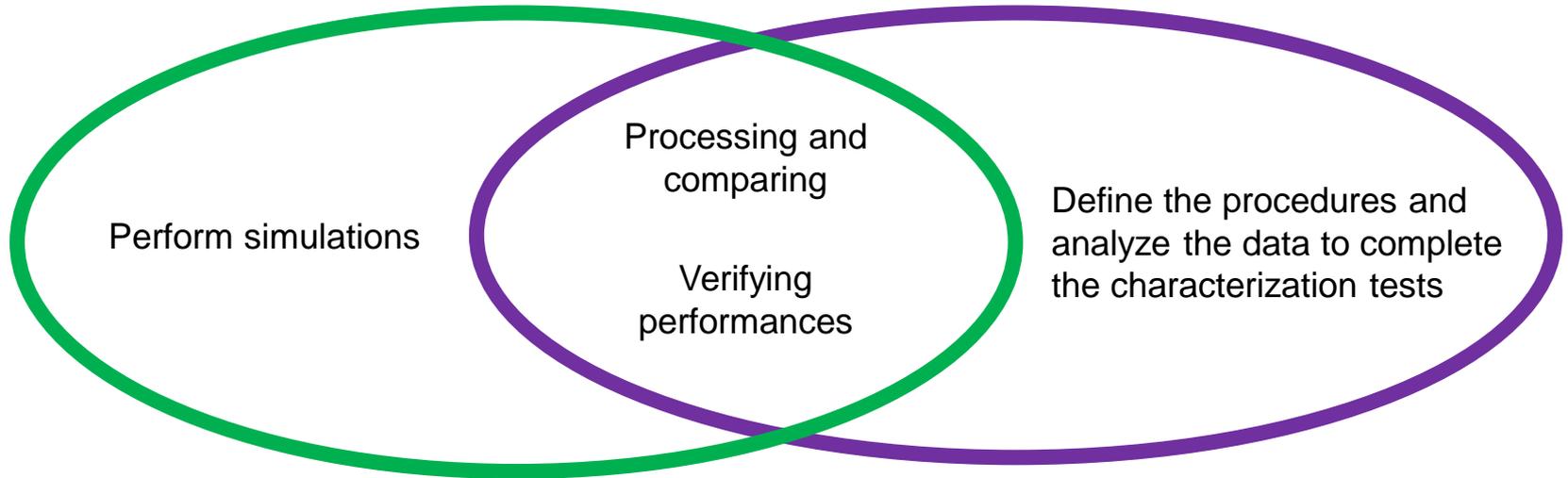
SIMBIO-SYS is composed by:

- **HRIC** (High Resolution Channel) the goal is the characterization of special surface targets with high resolution:
 - **400-900 nm** spectral range
 - **6m/px** Spatial resolution (at the best)
- **VIHI** (Visible Infrared Hyperspectral Imaging): the goal is to map the planet in order to provide the global mineralogical composition of the surface:
 - **400-2000 nm** spectral range.
 - **120 m/px – 480 m/px** spatial resolution
- **STC** (STereo Channel) is a double wide angle camera:
 - **410-930 nm** spectral range (2 panchromatics filters and 4 broad-band filters)
 - Total FoV **5.38° x 3.2°** (5.38° x 2.31° - 5.38° x 0.38°)
 - **58 m/px** spatial resolution (at best)

Calibration

Metis

STC



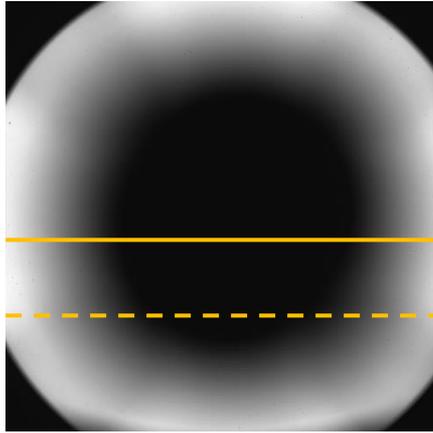
Perform simulations

Processing and
comparing

Verifying
performances

Define the procedures and
analyze the data to complete
the characterization tests

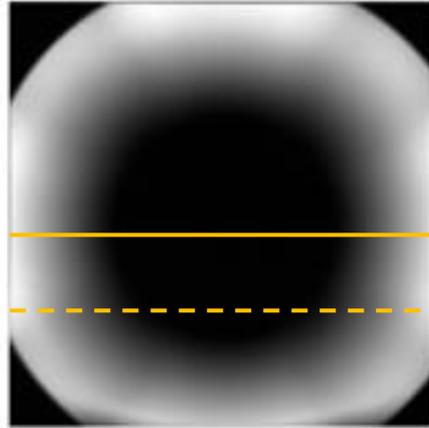
On-ground



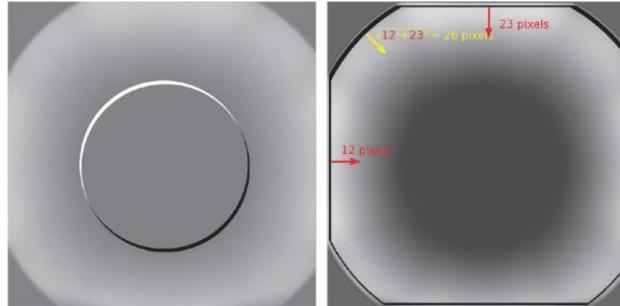
$t_{\text{exp}}=100\text{s}$



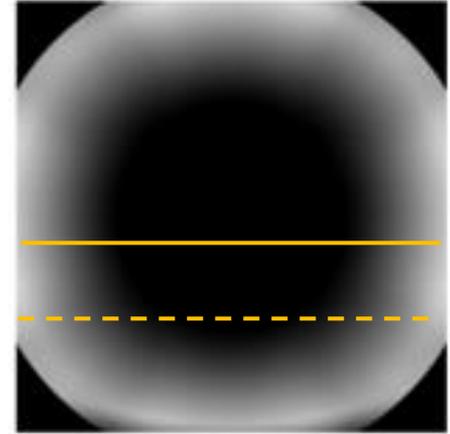
In-flight (reconstructed)



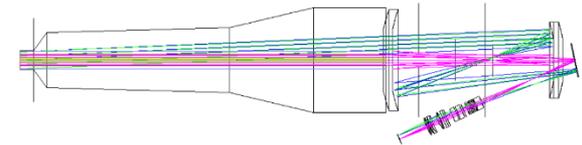
The centre of the image is changed because we move the IO to minimize the straylight



Simulated

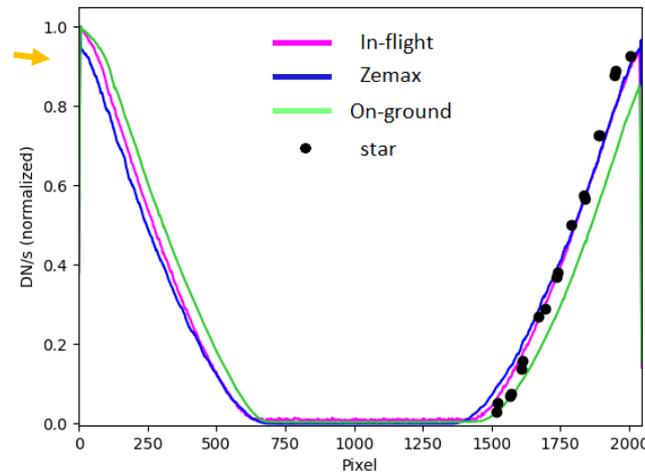
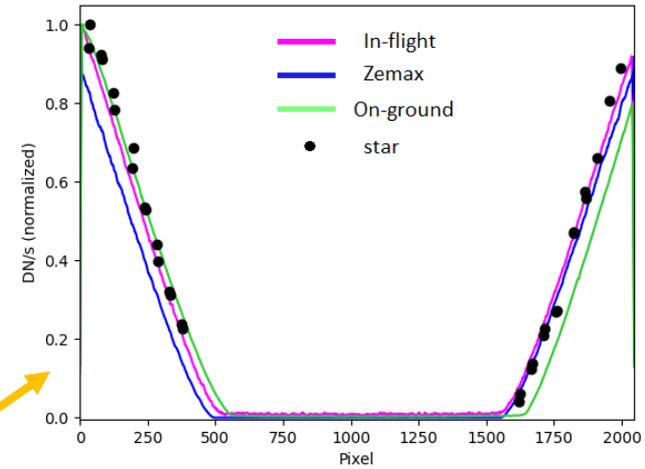
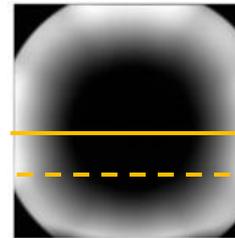
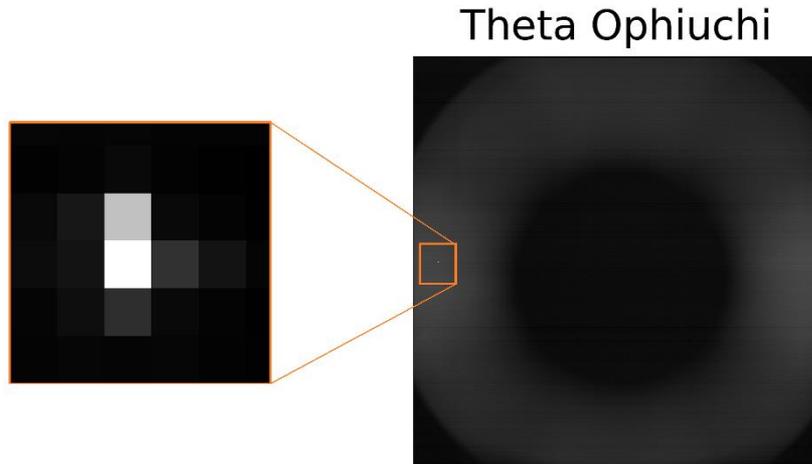


Raytracing simulation of Metis

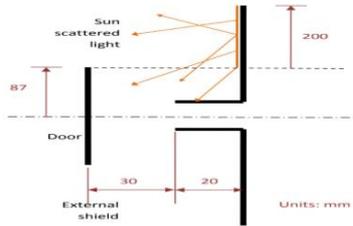


Theta Ophiuchi passed twice in front of Metis:

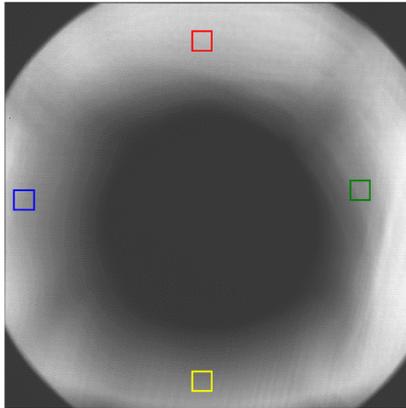
- STP 140 March 2021 (0,72 AU) - *row 1017*
- STP 182 December 2021 (1,01 AU) - *row 615*



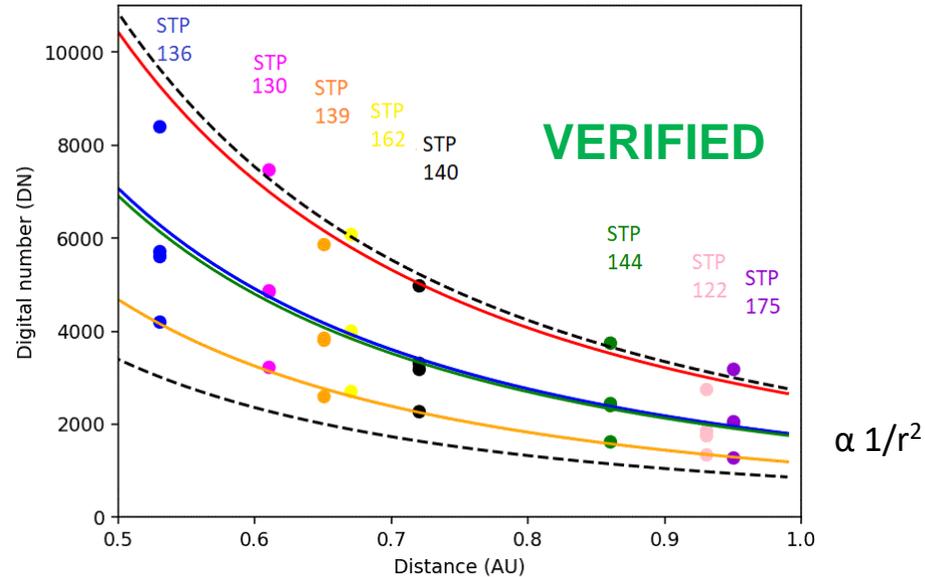
Verification of non degradation: 8 images at 8 different Astronomical distances



The light of the sun is reflected by the shield and, a portion is reflected towards the door.



Boxes of 100 x 100 pixels in the visible channel 4 fitting curves.



- we can use the retro-reflection of the door to estimate the optical elements possible degradation
- we can predict the effects of the possible degradation on the door images via ray tracing simulations .

Dark current acquisitions campaign:

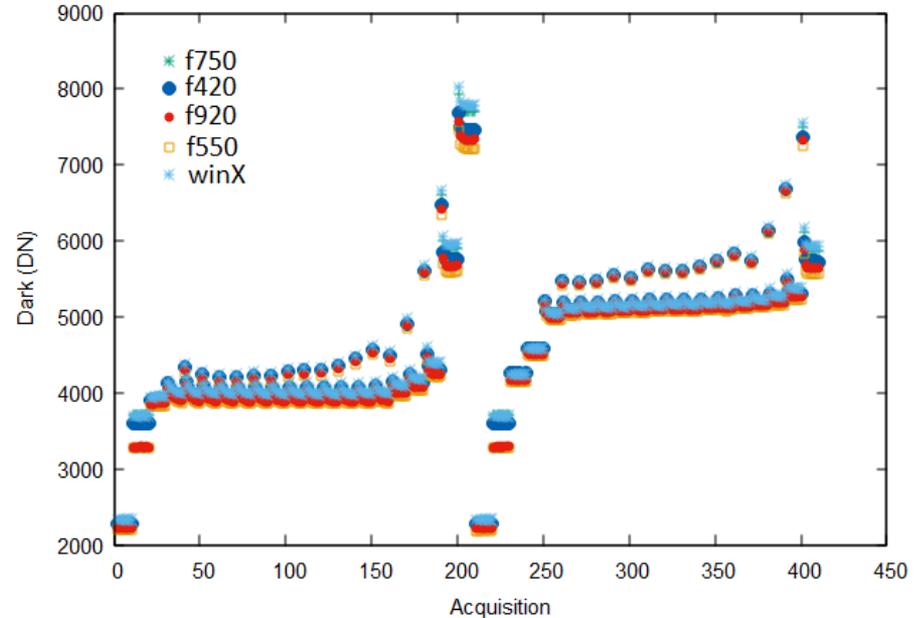
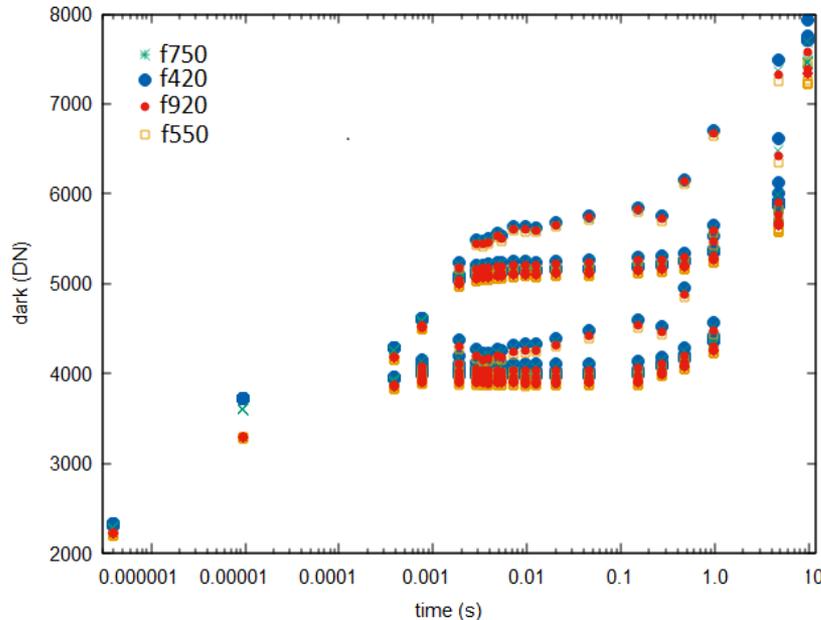
- the acquisition of a set of 10 dark-current images
- for a specific Integration Time (IT)
- with a specific Repetition Time (RT)

STC detector composed by:

- 2 Panchromatic (PanL and PanH)
- 4 colored filters (f750, f420, f920 and f550)

Start-End Rows		Vert Dim
2016	576px (strip 9) 1471px (strip 22)	
	F920	64px
1953 1808		
	F550	64px
1745 1610		
	PANL	384px
1227 820		
	PANH	384px
437 303		
192:319 (strips 3.4)	F420	64px
163 WinX 100 95		
	F750	64px
(0,0)	896px	

Mean value of the dark signal intensities acquired through the filter f750, f420, f920 and f550 at different exposure time.



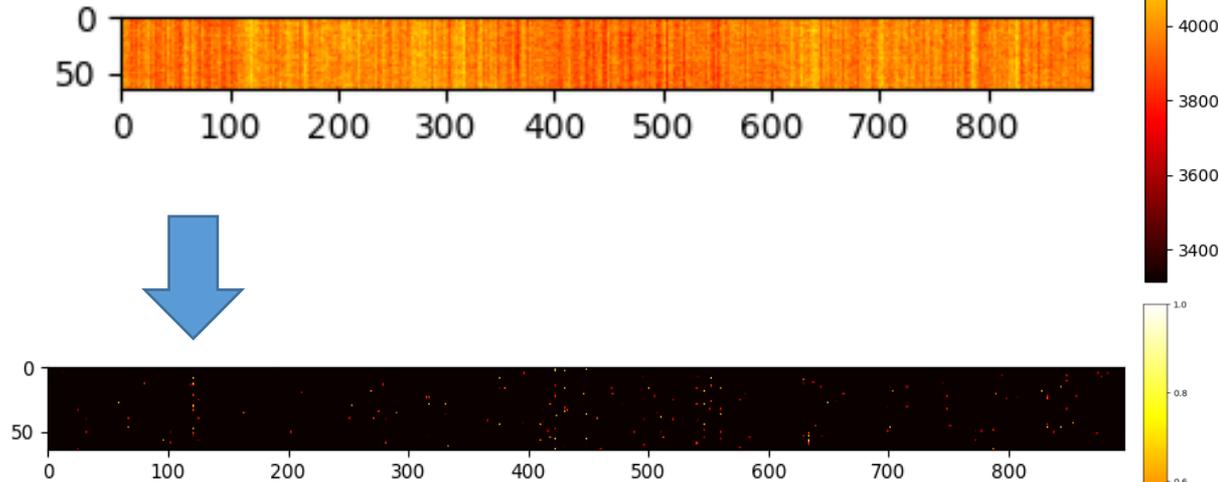
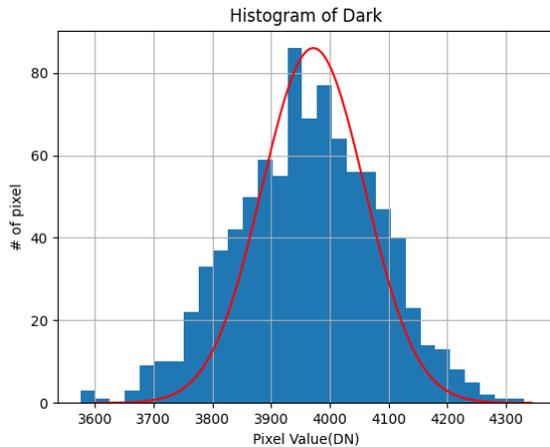
One possible solution is considering a small window named windows-x (WINX). It is a out-of-filter window, which is a region on the detector of dimension 64x128px sitting in the unilluminated part of the detector

VERIFIED PERFORMANCES (STC)

The thresholds for designating a pixel as “bad” were set to 3σ above or below the mean as appropriate.

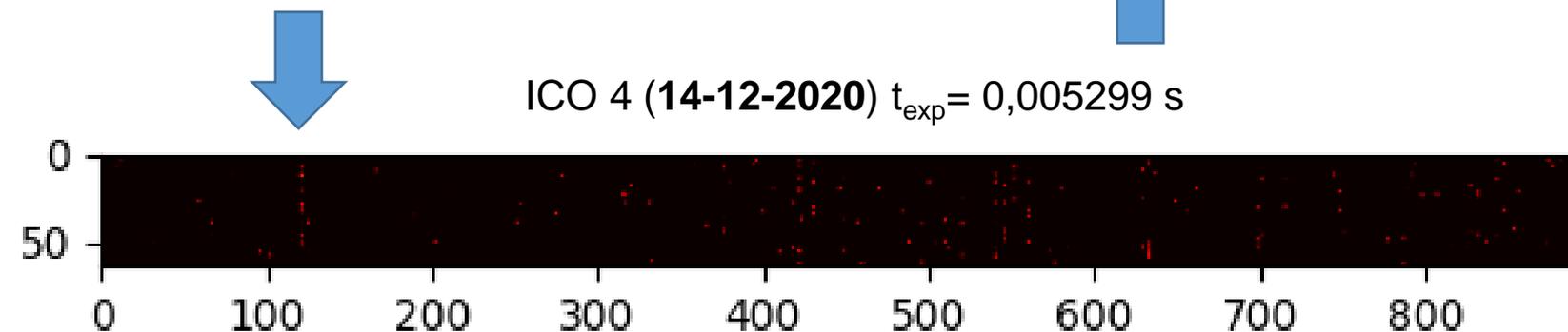
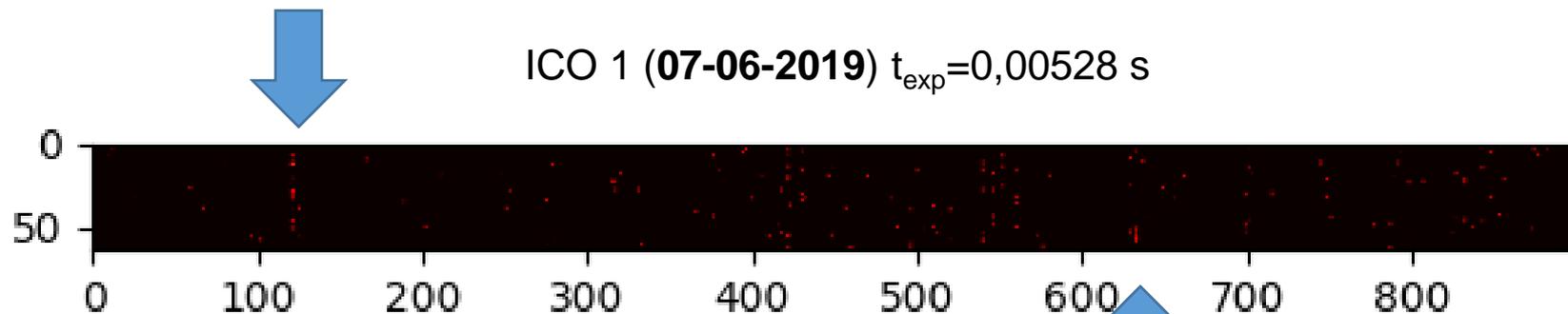
Therefore, I calculate for every image its histogram.

For Normally (Gaussian) distributed data only 0.3% of the pixels values would lie outside the mean $\pm 3\sigma$ range



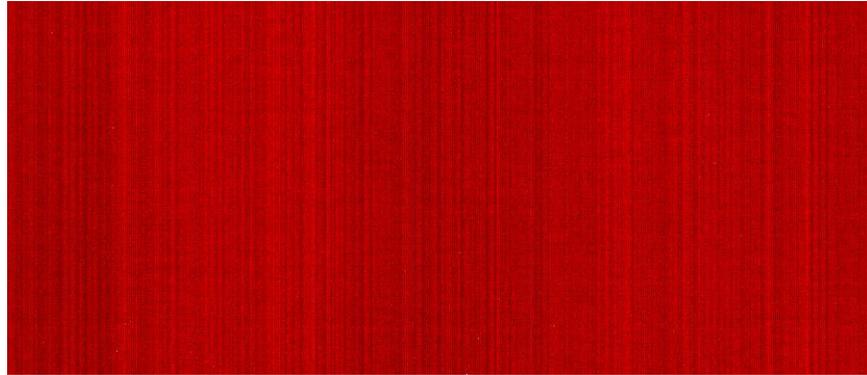
Take all value beyond **mean $\pm 3\sigma$**

- Lookup table (LUT): on each image analysis an increment of 1 on bad pixel coordinates
 - Divide for the 10 (number of the images)
- bad pixel on all images has the value of 1 on LUT (white)



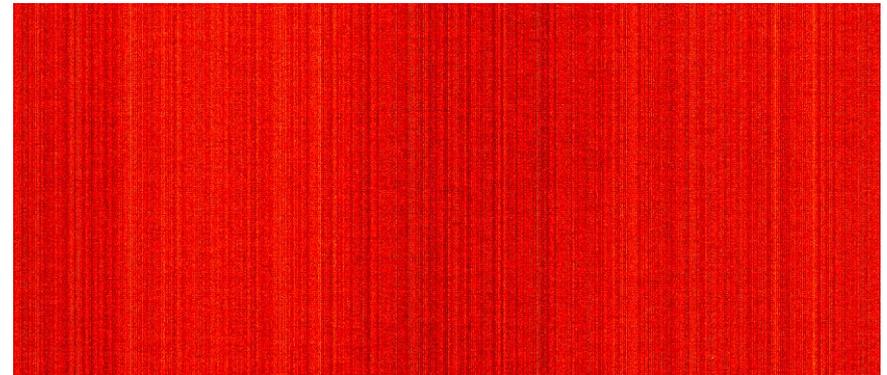
Similar column feature after 18 months -> readout noise.

Pan L



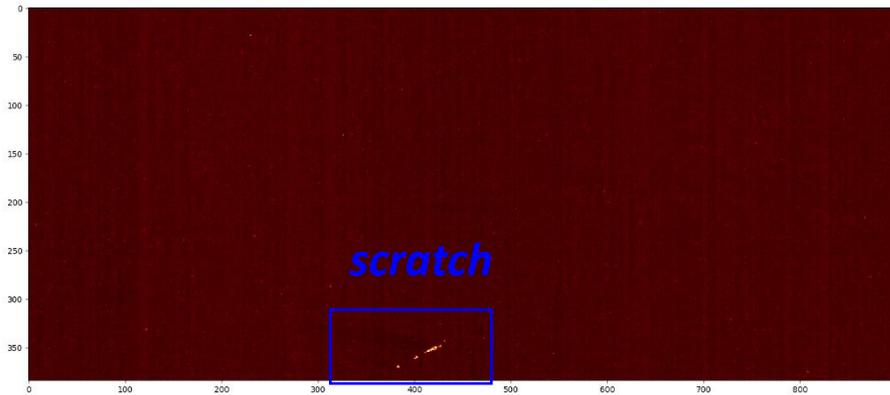
$4 \cdot 10^{-7}$ s

Panchromatic

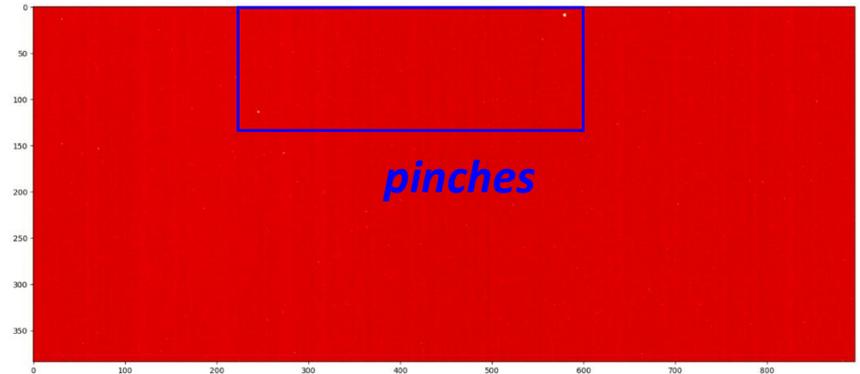


$4 \cdot 10^{-7}$ s

Pan H

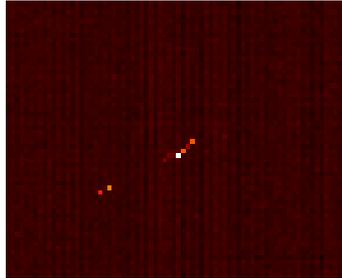
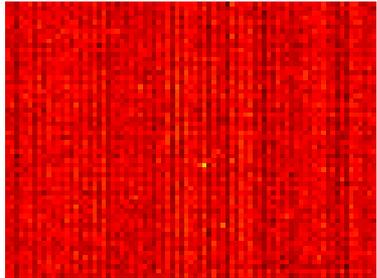


4,8 s



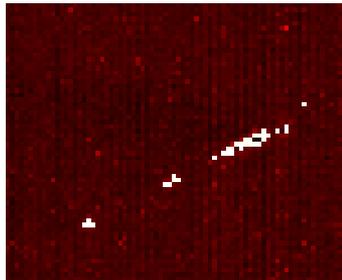
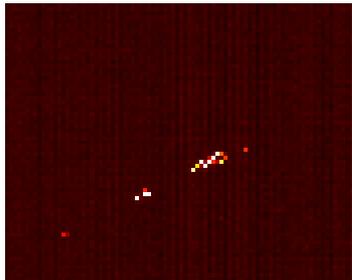
4,8 s

$5,76 \cdot 10^{-5}$ s **Pan L** 0,0024 s

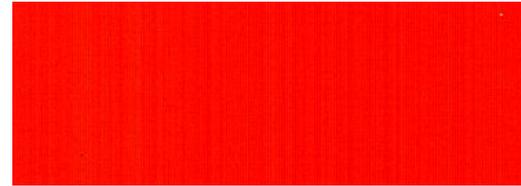


0,0096 s

4,8 s



0,0024 s **Pan H** 0,03 s

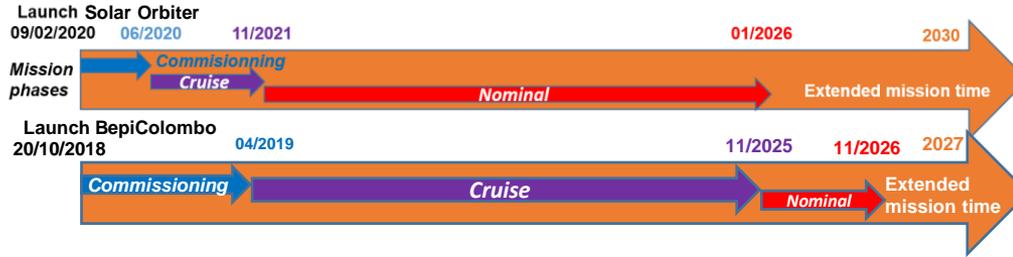


0,27 s

4,8 s



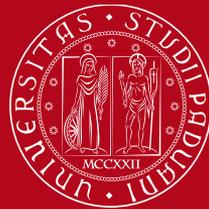
Acquisitions are now taken with the same intervals and times to understand if they are replicable and therefore correctable!



TASK TITLE	% OF TASK COMPLETE	T1				T2				T3				T4																								
		O	N	D	J	F	M	A	M	J	J	A	S	O	N	D	J	F	M	A	M	J	J	A	S	O	N	D	J	F	M	A	M	J	J	A	S	
Commissioning phase																																						
Review of state of the art of the Metis and STC	100%	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	
Calibration activities: comparison with on ground results	80%	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	
Optical performances	80%	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	
Cruise Phase																																						
Calibration sequence planning	70%	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	
Performances validation	75%	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█
Straylight	80%	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█
Data Analysis and results																																						
Calibration input for the pipeline	1%	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	
Support to pipeline implementation	1%	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█
Phd Thesis Writing																																						
...	1%	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	█	

Thanks for the attention

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